

Searches for parabottomonia in current and future B factories

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BNM 2008 Workshop, Atami 24-26/1/2007

Sunrise from Mt Fuji at Motosuko Lake,Yamanashi

Bottomonium

$b\bar{b}$ bound state
non-relativistic quark motion,
in perturbative regime

$$\alpha_S \sim 0.1$$

J,S,L are good quantum numbers

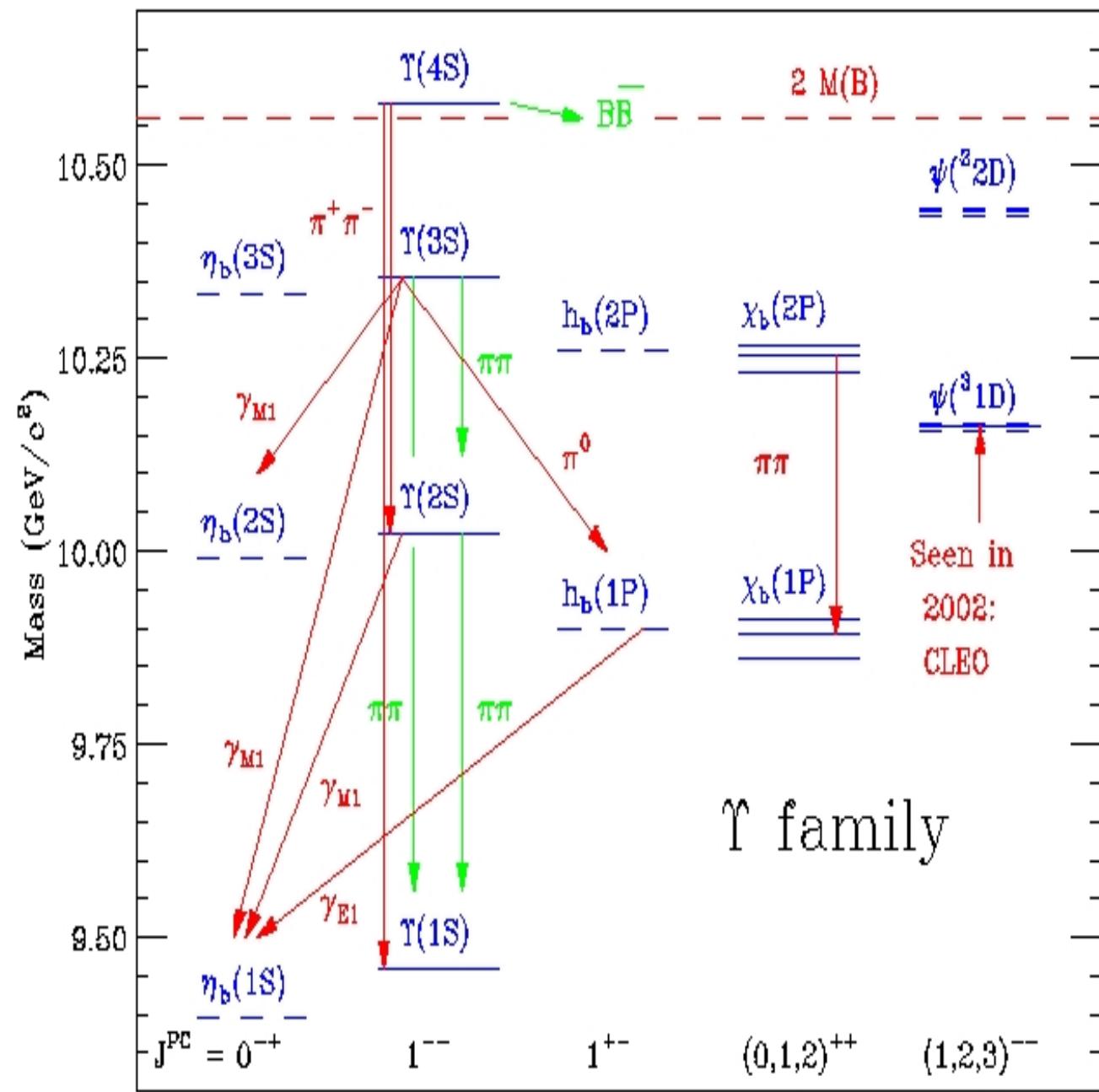
S=0 : *parabottomonium*
Not yet observed

S=1 : *ortobottomonium*
Y decay to 3 gluons
 $\Gamma[Y(1,2,3S)] = 53, 44, 26 \text{ keV}$

No 2S-2P Degeneration

Hyperfine Splitting : *unknown*

Fine Splitting : **15-30 MeV (1P)**



Quarkonium Working Group Activities



YELLOW REPORT : CERN-2005-005, ArXiv: hep-ph/0412158

QWG Workshops on Heavy Quarkonium:

QWG1: CERN, November 8 to 10, 2002

QWG2: Fermilab, September 20 to 22, 2003

QWG3: Beijing, October 12 to 15, 2004

QWG4: Brookhaven, June 27 to 30 , 2006

QWG5: DESY Hamburg, October 17 to 20, 2007

QWG6: NARA, 2-5 December 2008

Priority Number 1:

*FIND THE MISSING
BOTTOMONIA!!!*

Effective field theories (EFT)

- Great progress in understanding EFT's during last 15 years
NRQED (Caswell, Lepage) --> NRQCD(Bodwin,Braaten, Lepage)
- NRQCD, pNRQCD, vNRQCD: progress in calculation of perturbative terms
- Rigorous factorization of non-perturbative terms to be computed on the lattice
- Applicable to all aspects of the dynamics of heavy quarkonia:
spectroscopy, decays, production

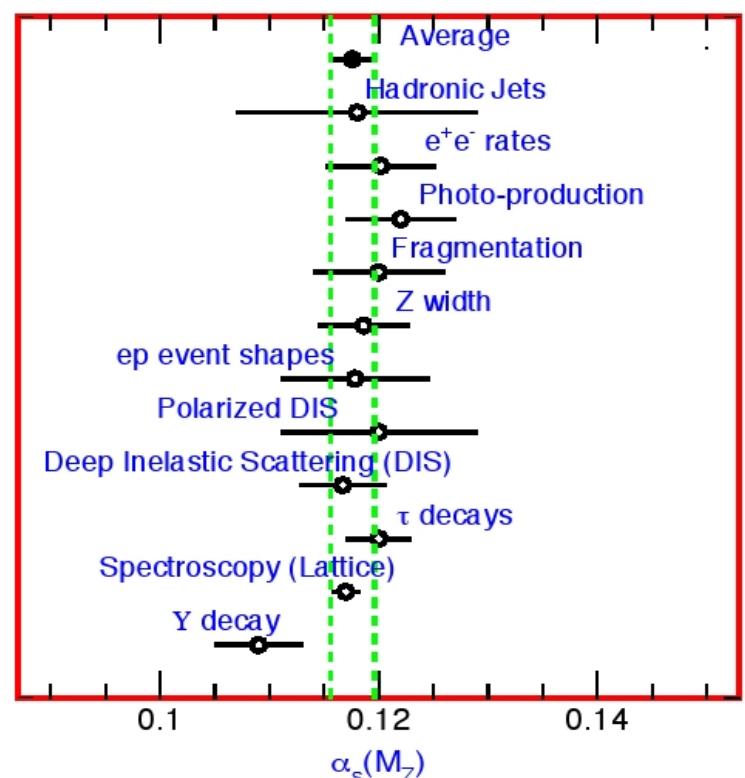
e.g.: Extraction of $\alpha_s(m_Y)$ from $Y \rightarrow \gamma gg / Y \rightarrow ggg$

$$\alpha_s(m_{Y_{1S}}) = 0.184 \pm 0.014 \rightarrow \alpha_s(m_Z) = 0.119 \pm 0.006$$

(PDG: 0.118 ± 0.003)

CLEO: PRD74(2006),012003

Brambilla et al: hep-ph/0702079



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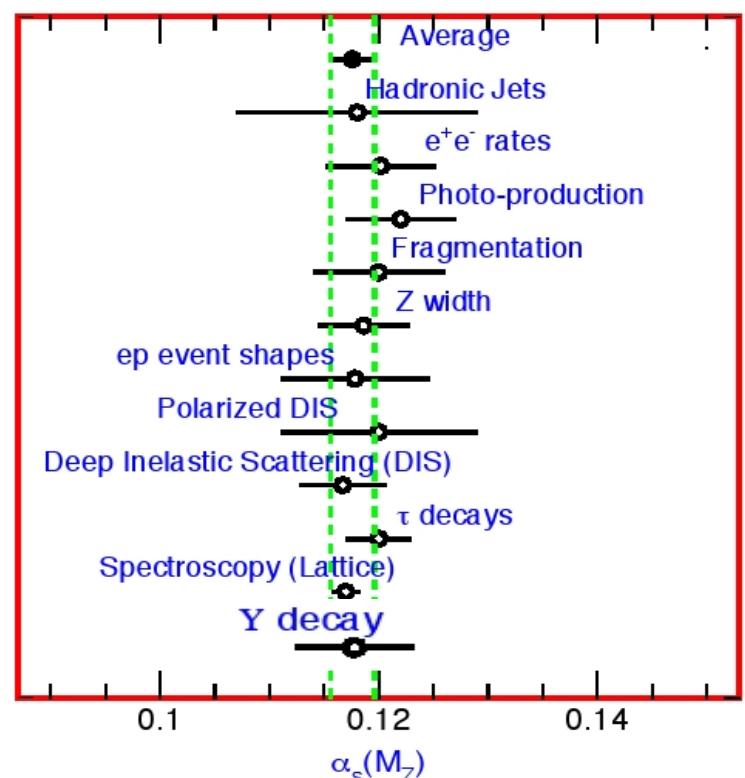
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NRQCD predictions

NLL calculation of $M(\eta_b)$:

$$M(\eta_b) = 9421 \pm 10(\text{th}) \pm 9(\delta\alpha_s) \text{ MeV}/c^2$$

Kniehl et al., PRL 92(2004), 242001

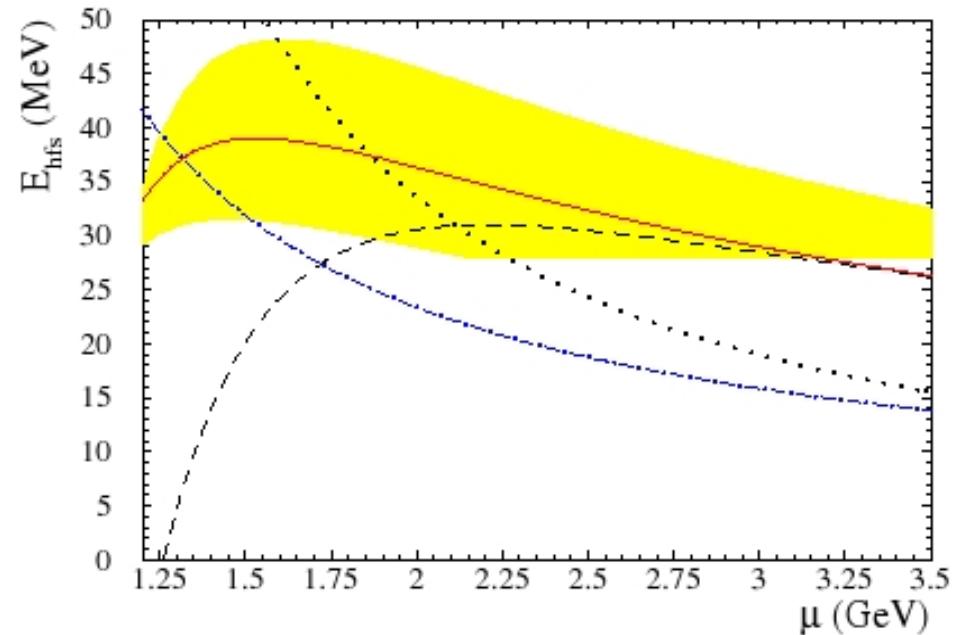


FIG. 1: HFS of 1S bottomonium as a function of the renormalization scale μ in the LO (dotted line), NLO (dashed line), LL (dot-dashed line), and NLL (solid line) approximations. For the NLL result, the band reflects the errors due to $\alpha_s(M_Z) = 0.118 \pm 0.003$.

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2-photon width at NNLL:

$$\Gamma(\eta_b(1S) \rightarrow \gamma\gamma) = 0.659 \pm 0.089(\text{th.})^{+0.019}_{-0.018}(\delta\alpha_s) \pm 0.015(\text{exp.}) \text{ keV}$$

Penin et al., NP B699 (2004),183

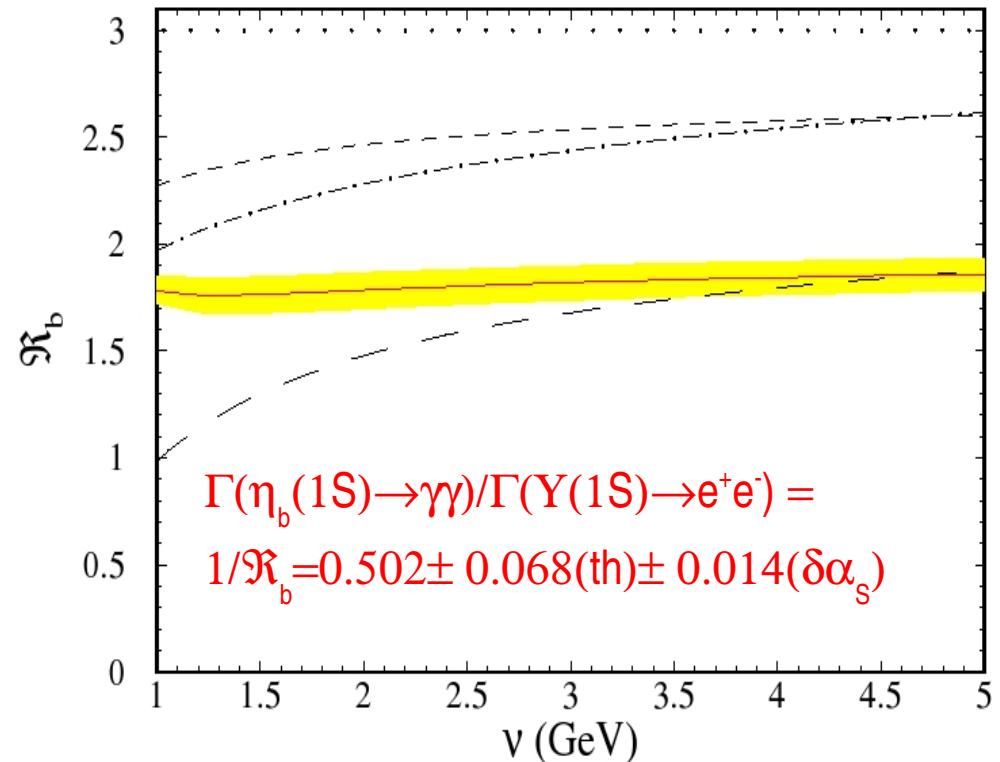


Figure 3: The spin ratio as the function of the renormalization scale ν in LO \equiv LL (dotted line), NLO (short-dashed line), NNLO (long-dashed line), NLL (dot-dashed line), and NNLL (solid line) approximation for the bottomonium ground state with $\nu_h = m_b$. For the NNLL result the band reflects the errors due to $\alpha_s(M_Z) = 0.118 \pm 0.003$.

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2-gluon width at NLO (for charmonium) :

$$\Gamma(\eta_c \rightarrow gg)/\Gamma(\eta_c \rightarrow \gamma\gamma) = 3.1 \pm 0.5(\text{th}) \pm 0.3(\delta\alpha_s) * 10^3 \quad \text{Exp}=3.6 \pm 1.2 * 10^3$$

Bodwin, Chen, PRD64(2001),114008; Bodwin, Petrelli, PRD66(2002),094011

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Kniehl et al., PRL 92(2004),242001

$$\Gamma(\eta_b) \sim 30 \text{ MeV} ??$$

$$\text{BR}(\eta_b \rightarrow \gamma\gamma) \sim 0.2 * 10^{-4} ??$$

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$$\Gamma(\eta_b(1S) \rightarrow \gamma\gamma) = 0.659 \pm 0.089(\text{th.})^{+0.019}_{-0.018}(\delta\alpha_s) \pm 0.015(\text{exp.}) \text{ keV}$$

Penin et al., NP B699 (2004),183

2-gluon width at NLO (for bottomonium) :

$$\Gamma(\eta_b \rightarrow gg)/\Gamma(\eta_b \rightarrow \gamma\gamma) = (Z_c/Z_b)^4 \quad \Gamma(\eta_c \rightarrow gg)/\Gamma(\eta_c \rightarrow \gamma\gamma) \sim 5 * 10^4$$

Bodwin, Chen, PRD64(2001),114008; Bodwin, Petrelli, PRD66(2002),094011

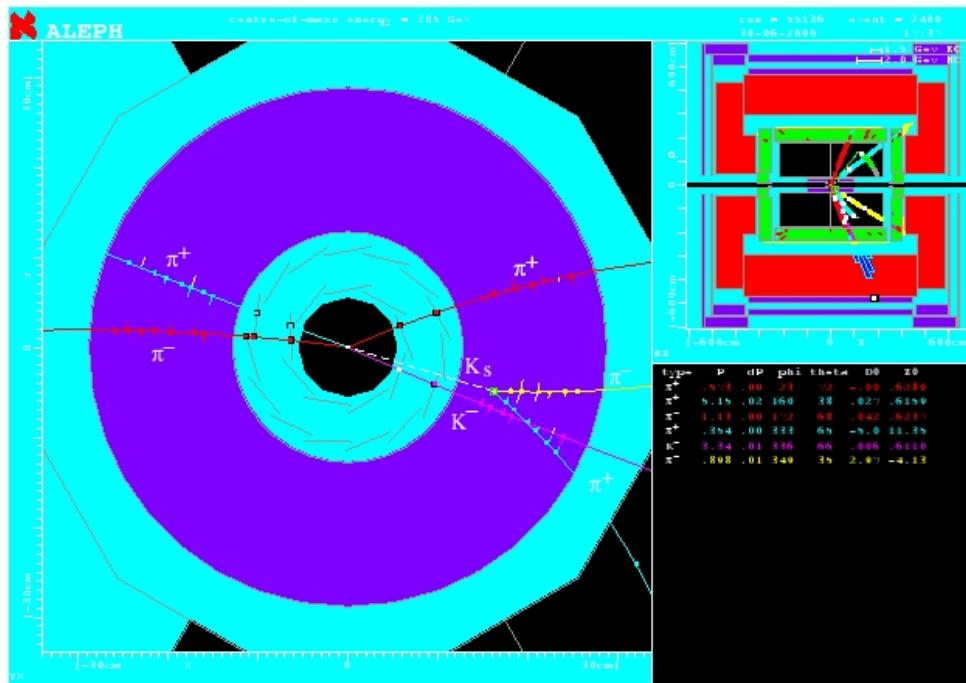
Searches for $\eta_b(1,2,3S)$ at LEP and Tevatron

Aleph, Delphi, L3 @ LEP

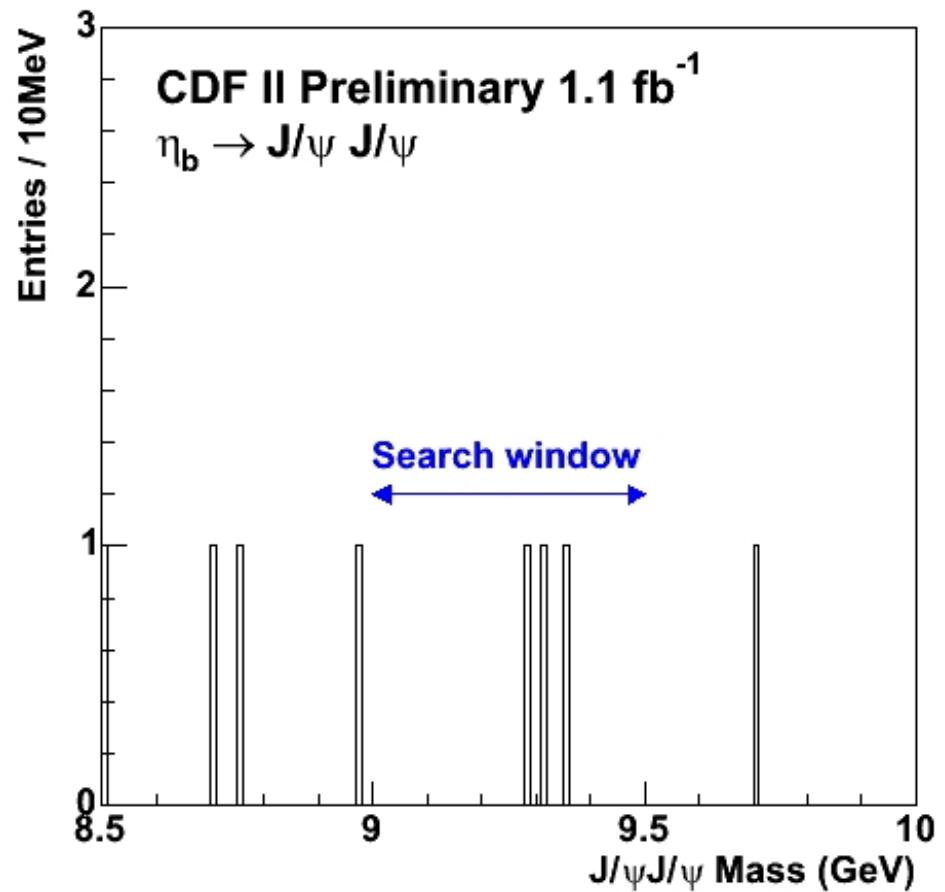
($\gamma\gamma \rightarrow \eta_b \rightarrow 4,6$ prongs)

few events, compatible with background

η_b Candidate Event



CDF @ Tevatron ($\eta_b \rightarrow J/\psi J/\psi$)



... but theory predicts $10^9 \eta_b / \text{fb}^{-1}$ ($= \mu\text{b}$)

[Braaten et al., PRD 63, 094006 (2001)]

mass $m = 9.30 \pm 0.02 \pm 0.02$ GeV

Inclusive searches in $\Upsilon(nS)$ decays

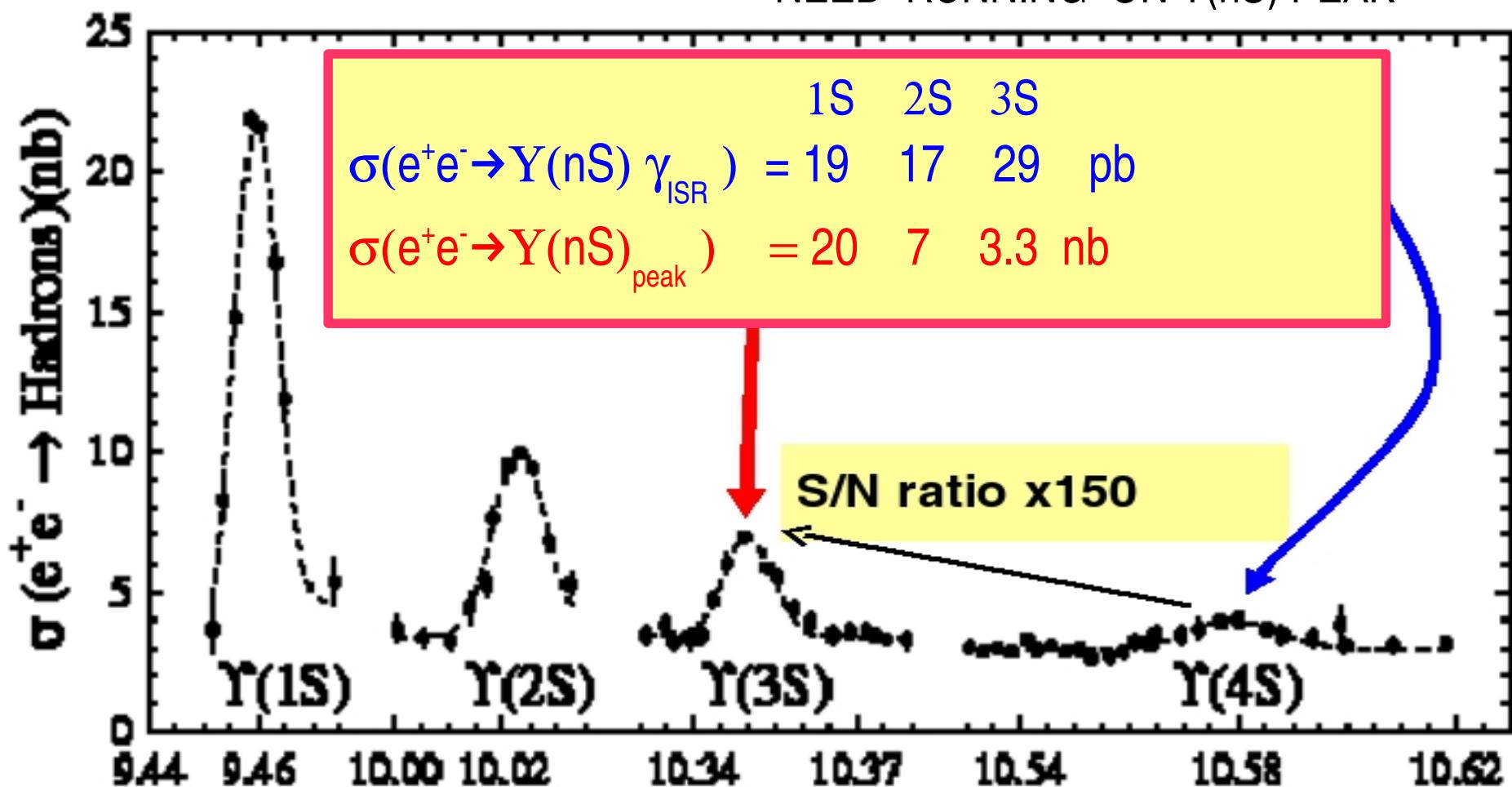
PRO's

- Do not pay the penalty of small exclusive Branching Ratios.
- If unbiased, allow to determine absolute BR's

CON's : ISR samples are unusable:

- w/ tagged ISR photon ($\sim 0.9 \text{ M in } 1 \text{ ab}^{-1}$) : limited statistics
- w/ untagged ISR photon ($\sim 28 \text{ M in } 1 \text{ ab}^{-1}$) : smearing in resolution

NEED RUNNING ON $\Upsilon(nS)$ PEAK



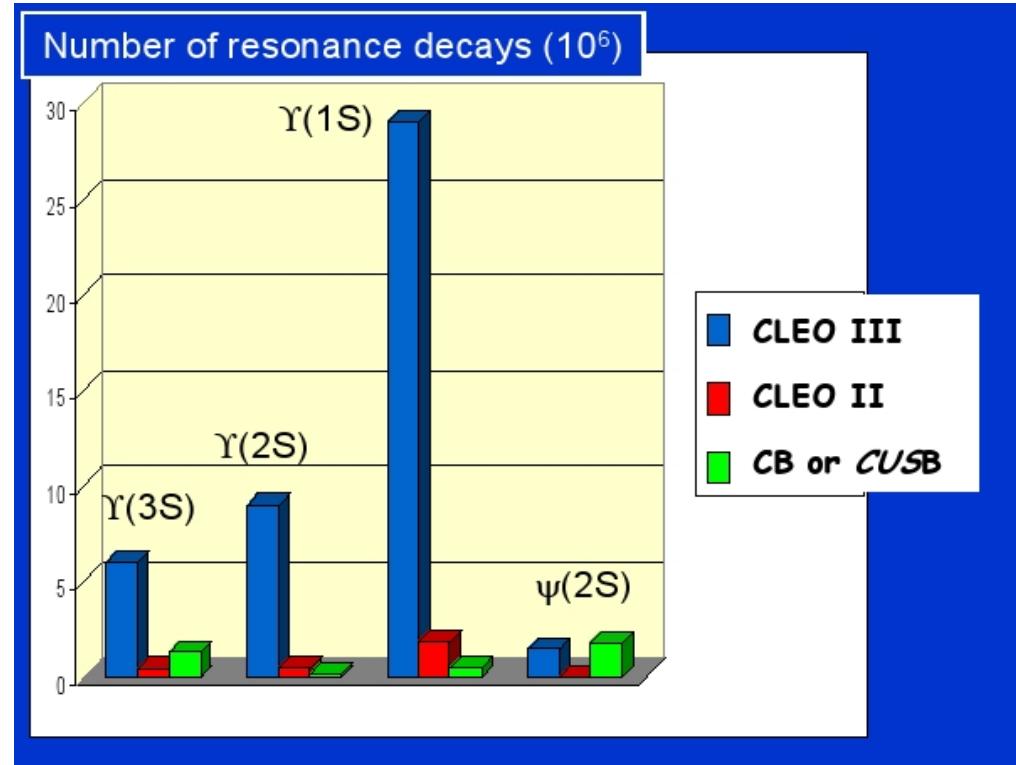
Data Samples in B factories

CLEO-III (2003):

$1.06 \text{ fb}^{-1} \Rightarrow 21 \text{M } Y(1S)$

$1.31 \text{ fb}^{-1} \Rightarrow 9.3 \text{M } Y(2S)$

$1.39 \text{ fb}^{-1} \Rightarrow 5.9 \text{M } Y(3S)$



BELLE (2006):

$2.9 \text{ fb}^{-1} \Rightarrow 11 \text{ M } Y(3S)$

BaBar (now!)
goal: $30 \text{ fb}^{-1} \Rightarrow 100 \text{ M } Y(3S)$

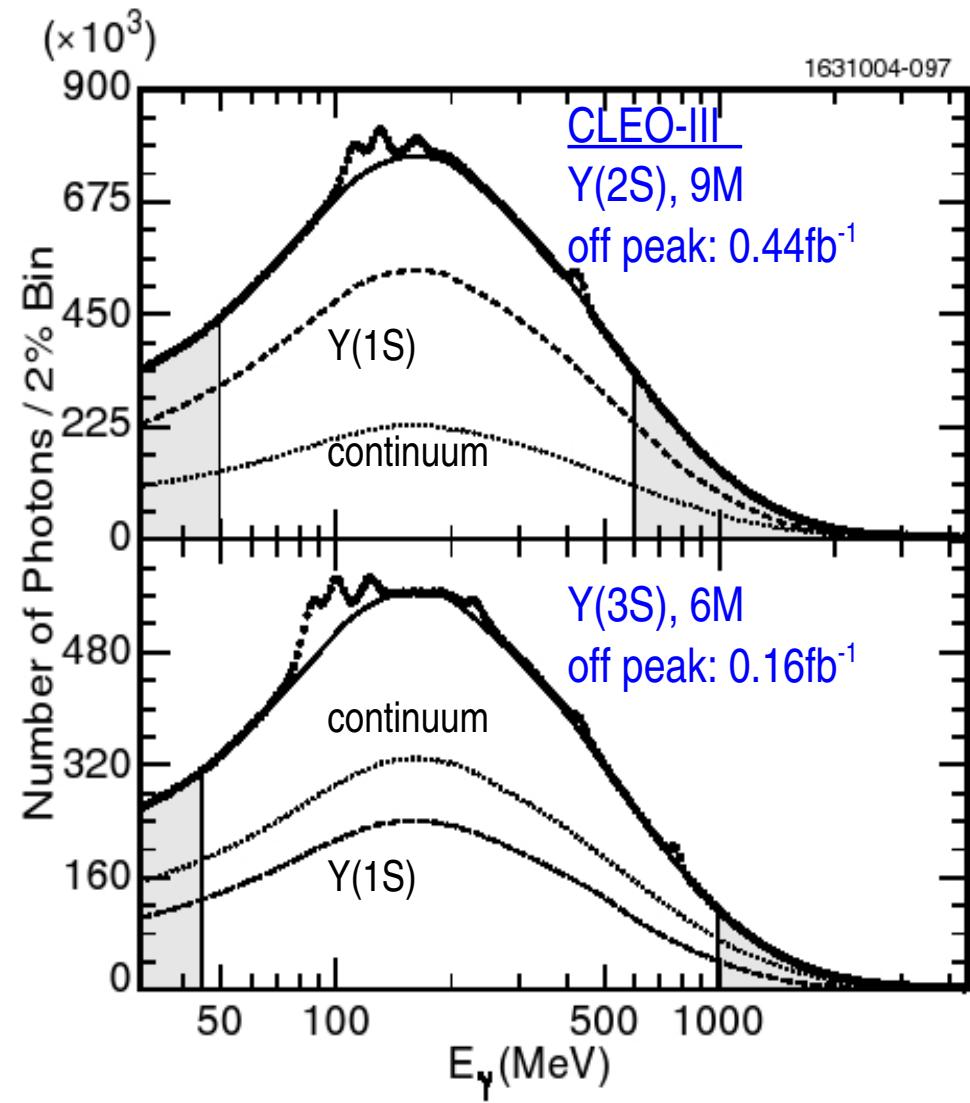
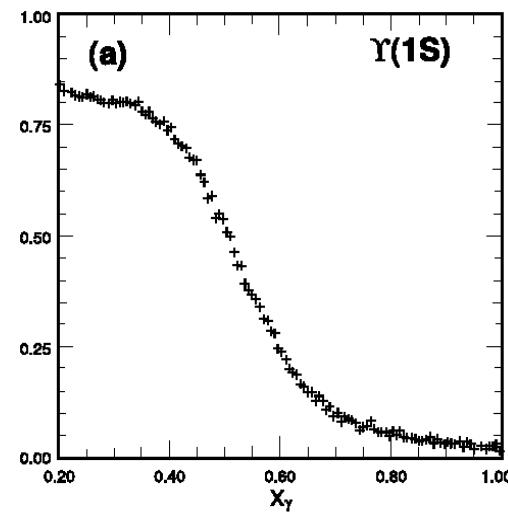
The challenge

Inclusive photon spectra at $Y(nS)$ are populated by comparable amounts by photons from transitions to other charmonia, and annihilations to LH.

On average about 7-8 photons per hadronic event are observed, a large fraction from decays of π, η, ω and other light hadrons.

The visible peaks are from known single photon transitions to χ_b 's at 10% level or cascades at 1% level.

(see CLEO:
ArXiv:0704.2766)



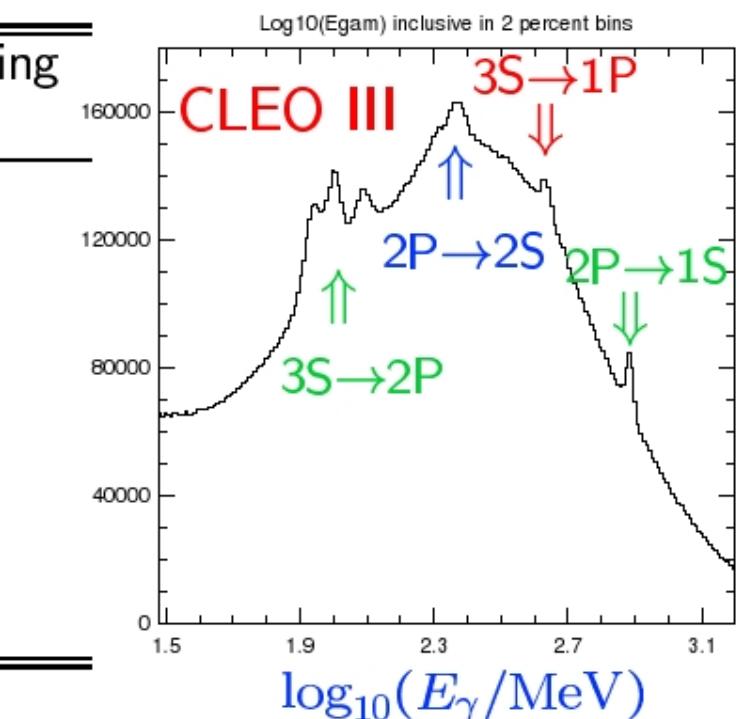
Inclusive searches w/ current data

Expected significances with
current CLEO samples
(assuming BR=10⁻⁴ and Gamma=10MeV)

	Emin-Emax (MeV)	FWHM (2%bins)	Bkg/ sample	Signal/ sqrt(Bkg)
3S->3S	30 100	10	3M/6M	0.6k/1.7k
3S->2S	330 426	3	1.1M/6M	0.6k/1k
3S->1S	856 947	1	160k/6M	0.6k/0.4k
2S->2S	30 100	10	4M/9M	0.9k/2k
2S->1S	544 638	2	0.55M/9M	0.9k/0.75k
1S->1S	30 100	10	10M/20M	2k/3k

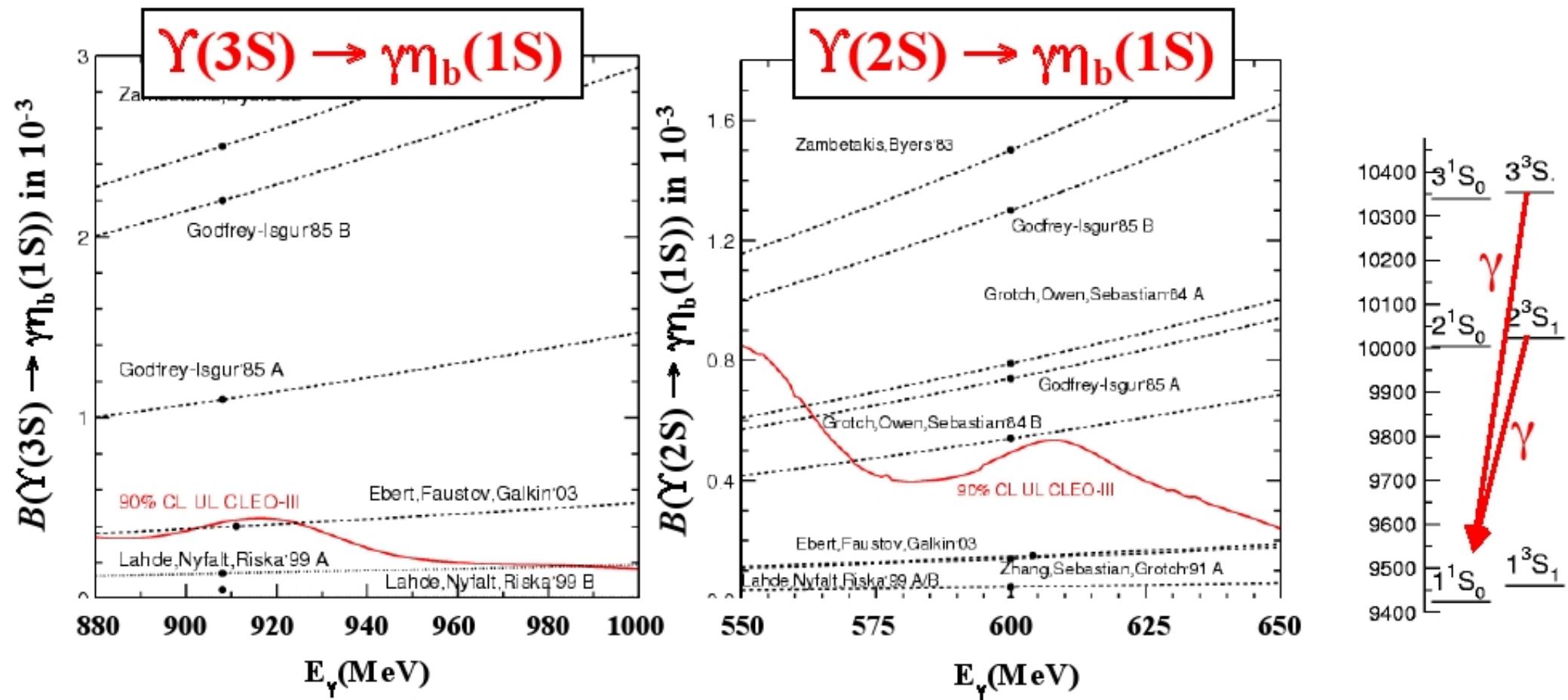
Photons above 400 MeV in 5.88 M $\Upsilon(3S)$ (Kwong + JLR, 1988, updated):

Energy (MeV)	Transition	Number of events (K)	Doppler broadening %	MeV
422	$\chi_{b1} \rightarrow \Upsilon(1S)$	32.3	4.7	20
433	$\Upsilon(3S) \rightarrow \chi_{b2}$	63.4	0	0
443	$\chi_{b2} \rightarrow \Upsilon(1S)$	24.3	4.5	20
453	$\Upsilon(3S) \rightarrow \chi_{b1}$	43.2	0	0
484	$\Upsilon(3S) \rightarrow \chi_{b0}$	17.6	0	0
741	$\chi'_{b0} \rightarrow \Upsilon(1S)$	2.1	1.2	8.9
765	$\chi'_{b1} \rightarrow \Upsilon(1S)$	75.3	1.0	7.5
777	$\chi'_{b2} \rightarrow \Upsilon(1S)$	45.6	0.8	6.4

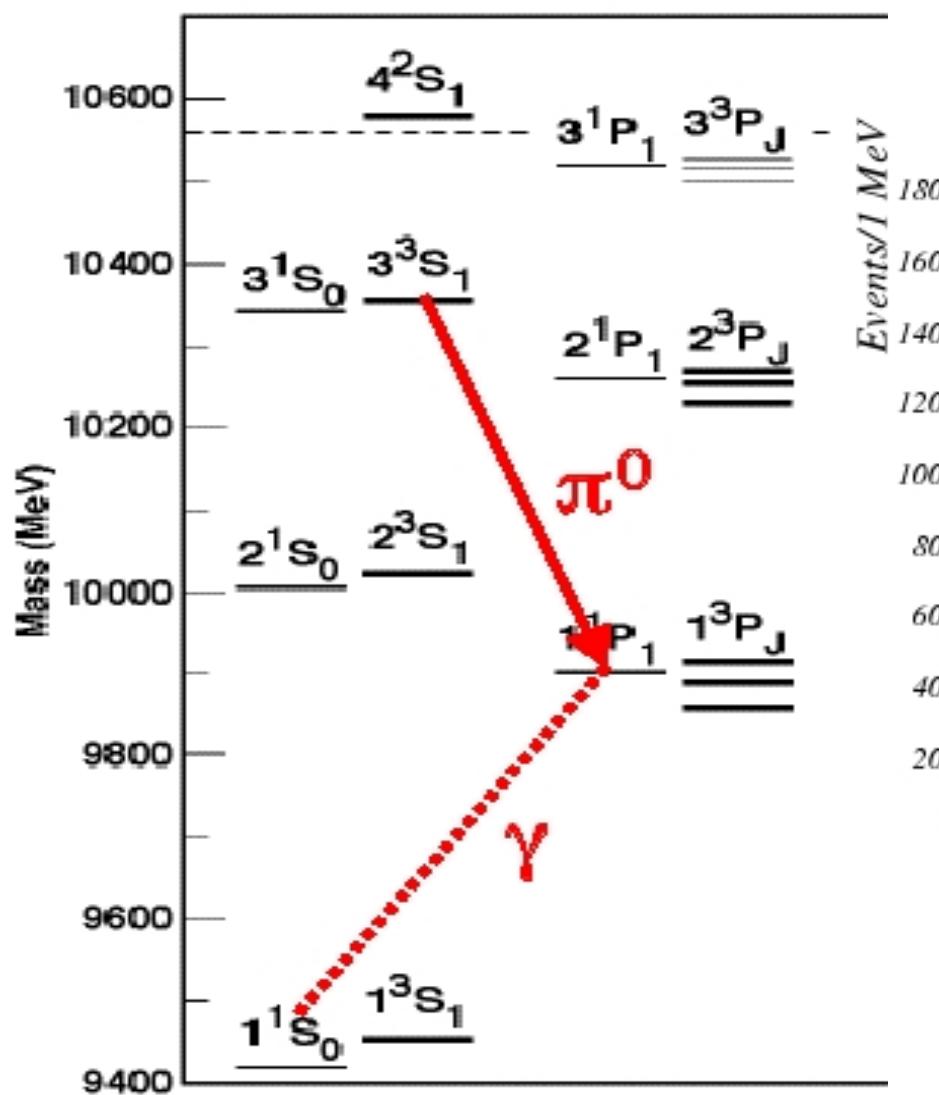


Inclusive searches of $\eta_b(1\text{S})$ at CLEO

- **Hindered M1 transition:** $\Gamma_{M1} \propto \frac{e_Q^2}{m_Q^2} \left| \langle nL | n'L \rangle \right|^2 E_\gamma^3$
- But $E_\gamma \sim 911$ (**604**) MeV from $\Upsilon(3\text{S})$ ($\Upsilon(2\text{S})$) $\rightarrow \gamma \eta_b(1\text{S})$ with $M(\eta_b) \sim 9400$ MeV/c².
- CLEO has already set ULs (90%CL) on these *BR*'s (PRL94,032001)

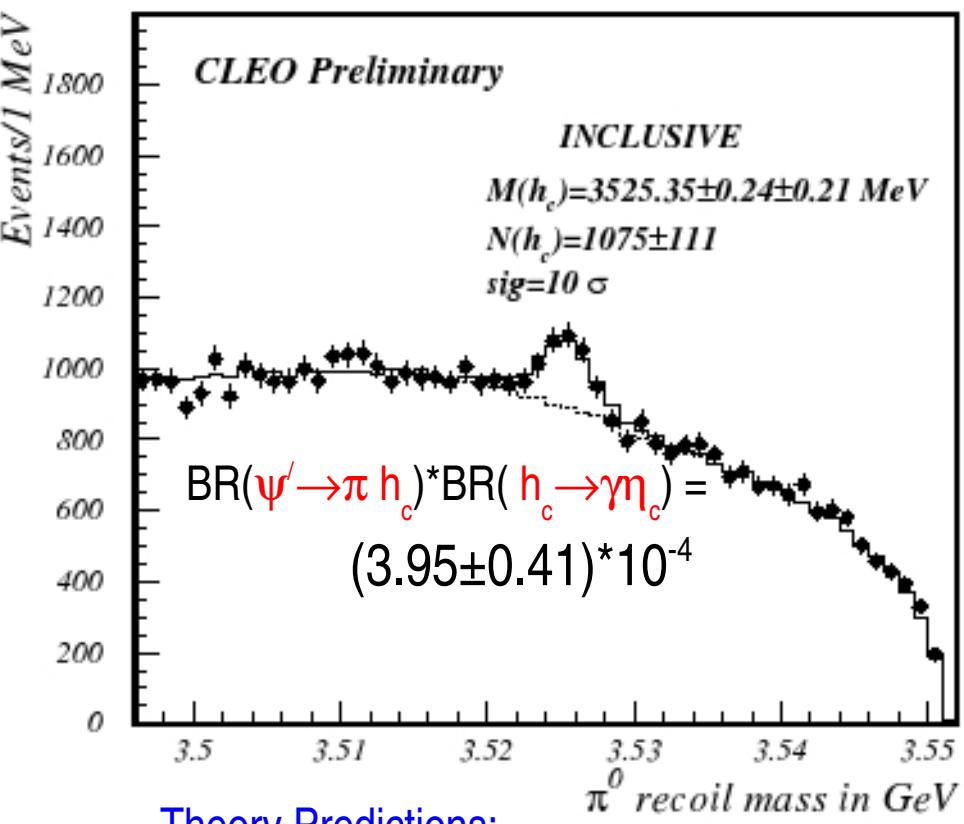


Double hunt: h_b (1P) and η_b (1S) at CLEO (and Belle)



Search for h_b (1P) in Bottomonium

Discovery of h_c (1P) in Charmonium :



Theory Predictions:

Voloshin: Sov.JNP43 (1986), 1011

Godfrey, Rosner, PRD 66 (2002), 014012

$\text{BR}(\psi' \rightarrow \pi h_b)^* \text{BR}(h_b \rightarrow \gamma \eta_b) = 4 * 10^{-4}$

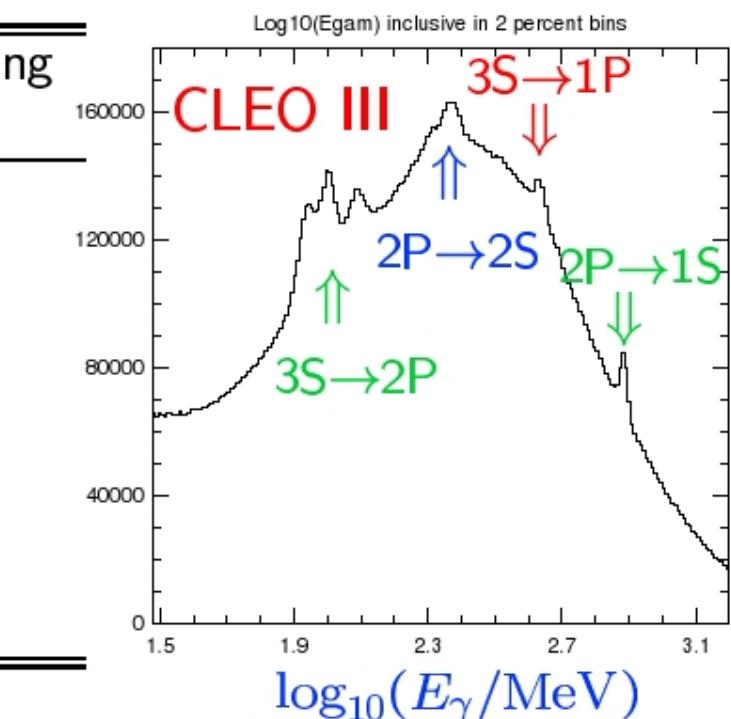
Inclusive searches w/ 100MY's

Expected significances with
future BaBar Y(3S) sample and with
comparable ones on Y(1,2S)
(assuming BR=10⁻⁴ and Gamma=10MeV)

	Emin-Emax (MeV)	FWHM (2%bins)	Bkg	Signal/ sqrt(Bkg)
3S->3S	30 100	10	50M	10k/7k
3S->2S	330 426	3	17M	10k/4k
3S->1S	856 947	1	2.5M	10k/1.6k
2S->2S	30 100	10	44M	10k/6.5k
2S->1S	544 638	2	6M	10k/2.5k
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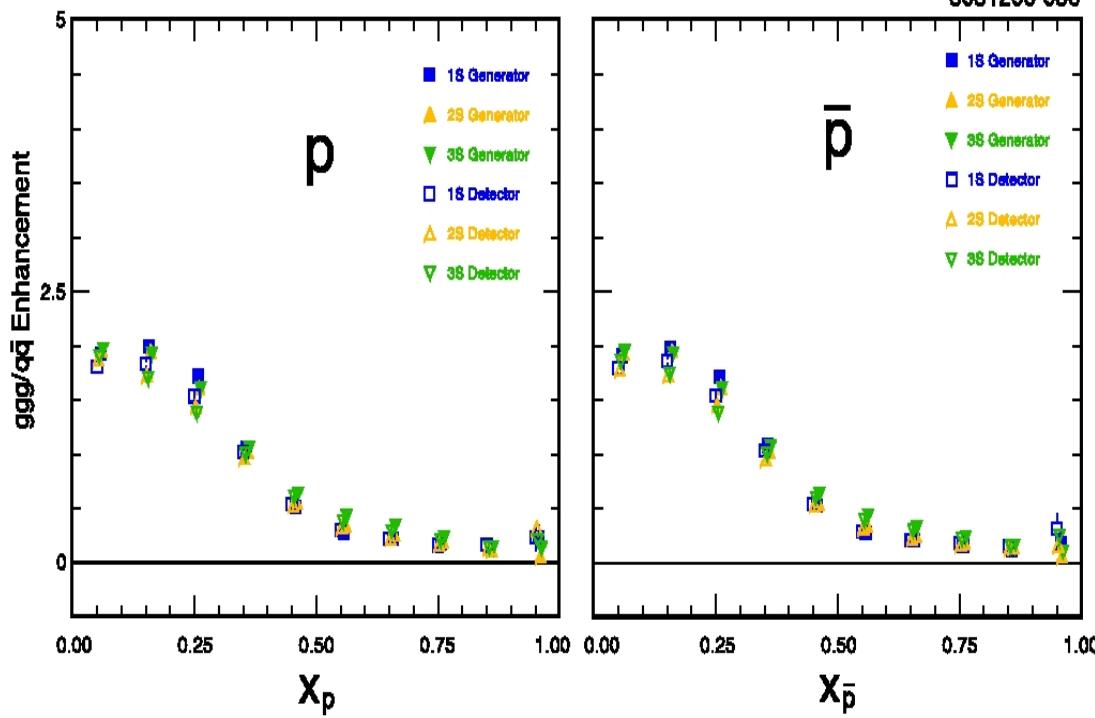
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Semi-inclusive signatures ?

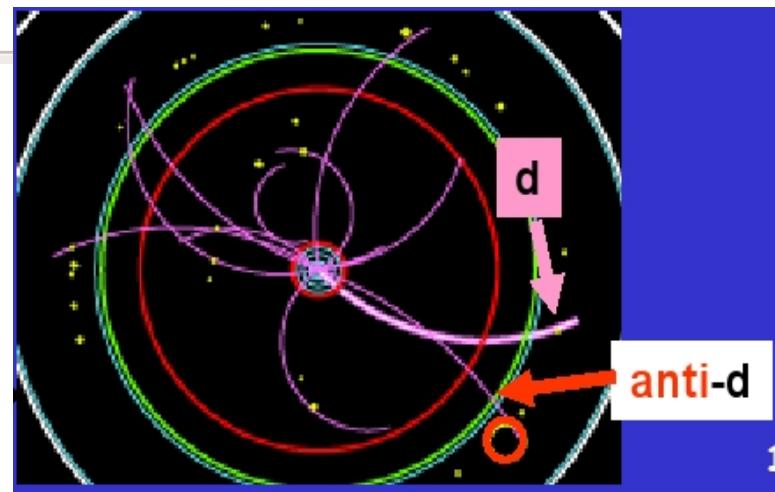
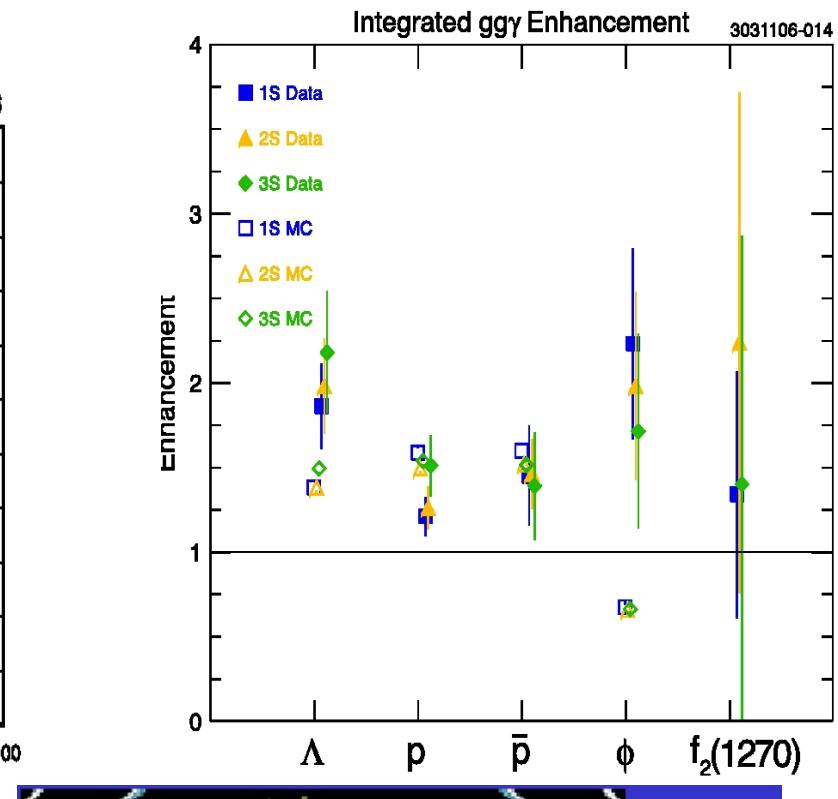
ArXiv:0704.2766

Enhanced production of low P protons, antiprotons,
 Λ 's in $Y(1,2,3\text{ S})$ decays if compared to continuum



... but $BR(p+)$ at few percent level : too low with current samples, but may help with 100M Y 's.

Anti-deuteron production observed in $Y(1S)$ and $Y(2S)$ decays ($BR=3.3\pm0.5 \cdot 10^{-5}$) and not in continuum ($BR<10^{-5}$) [Cleo: [PRD 75, 12009\(2007\)](#)]



Parabottomonia: other pathways

$$Y(3S) \rightarrow \pi\pi h_b(1P) \rightarrow \pi\pi \gamma\eta_b(1S)$$

Old CLEO-II limit at 0.18% with 0.46 M $Y(3S)$ sample

Expect limits down to 0.05 % with the current sample.

Need 20 times more statistics to get to theory predictions :

Voloshin (Sov.J.Nucl.Phys. 43,1011) $< 10^{-4}$

Kuang (hep-ph/0601044) $\sim 10^{-4}$

$$Y(3S) \rightarrow \gamma\chi_{b0}(2P) \rightarrow \gamma \eta\eta_b(1S)$$

Voloshin (Mod.Phys.Lett.:A19,2985(2004)):

Prediction for $\chi_{b0}(2P) \rightarrow \eta\eta_b(1S)$: $\sim 10^{-3}$

Recent observation (5σ) of $Y(2S) \rightarrow \eta Y(1S)$:

preliminary CLEO result : 2.5×10^{-4}

Summary

Study of narrow Υ decays is an extremely active field of research, both for our understanding of QCD on bound states and for searches of new physics.

The 5 states of parabottomonium expected from theory below open beauty threshold are still unfound. With less than 4 fb^{-1} taken in 2003, CLEO is still leading the research on these states, and will hopefully give us first hints of their existence. But a systematic program of studies on the 'dark side of parabottomonium' will need at least 100x more statistics

Even only 1% of the 50 ab^{-1} expected from super B factories, if taken at the peak of $\Upsilon(nS)$ states, will allow to study the holy grail of heavy quarkonium spectroscopy



Further studies with ~100 M $Y(nS)$ samples

Studies on χ_{bJ} (1,2P) states : **exclusive** decays to light hadrons, $\gamma\gamma$ widths

$Y(3S) \rightarrow \gamma \chi_b$ (2P) : 13M,12M,6M χ_{bJ} (2P) with $J=2,1,0$

$Y(2S) \rightarrow \gamma \chi_b$ (1P): 7M,7M,4M χ_{bJ} (1P) with $J=2,1,0$

Studies on hindered radiative transitions allow to quantify the (small) relativistic effects in bottomonium

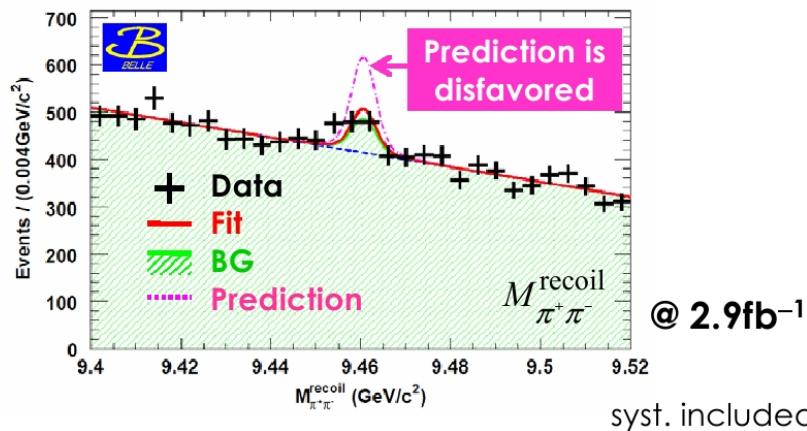
$Y(3S) \rightarrow \gamma \chi_b$ (2P) $\rightarrow \gamma\gamma Y(1S) \rightarrow \gamma\gamma\mu\mu$: 25k,28k,7k via χ_{bJ} (2P) with $J=2,1,0$

$Y(3S) \rightarrow \gamma \chi_b$ (1P) $\rightarrow \gamma\gamma Y(1S) \rightarrow \gamma\gamma\mu\mu$: 4k,6k, ~0.5k via χ_{bJ} (1P) with $J=2,1,0$

Dark matter searches in $\Upsilon(1S)$ decays

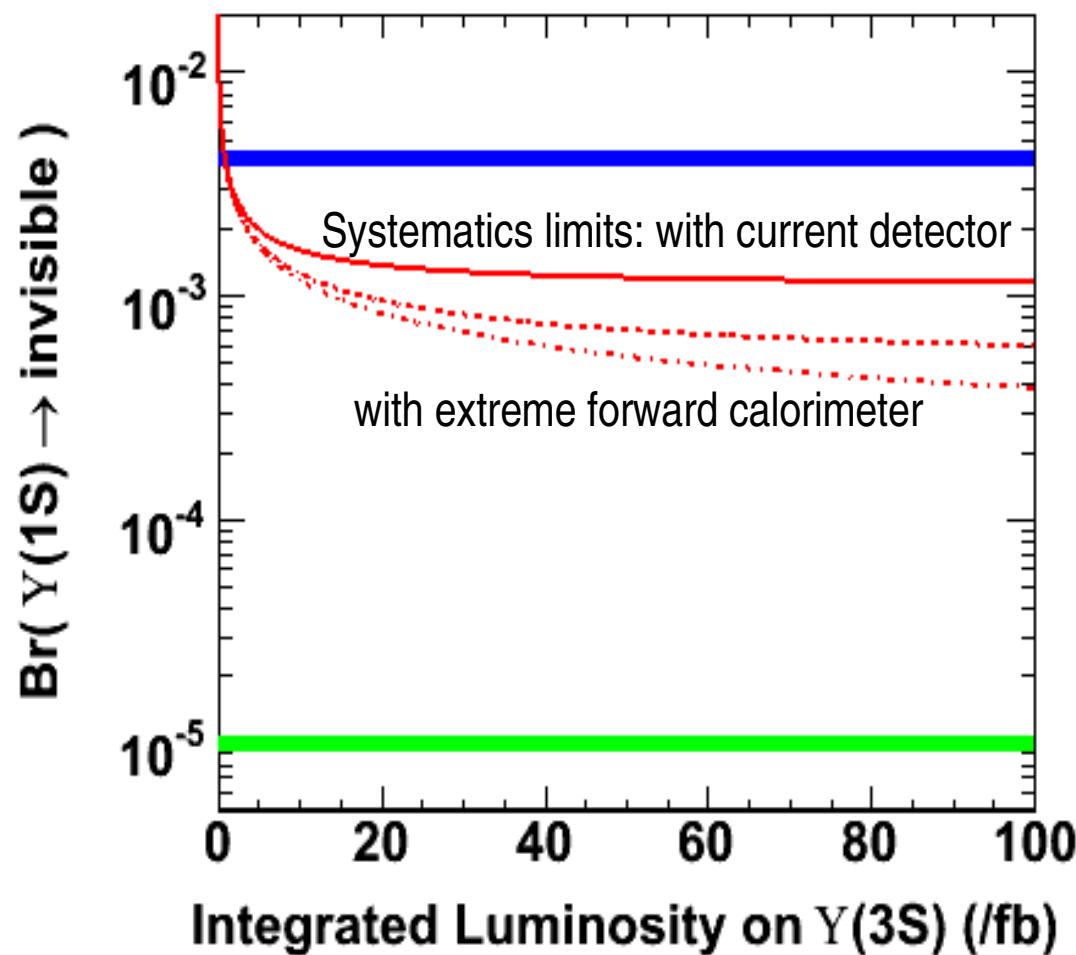
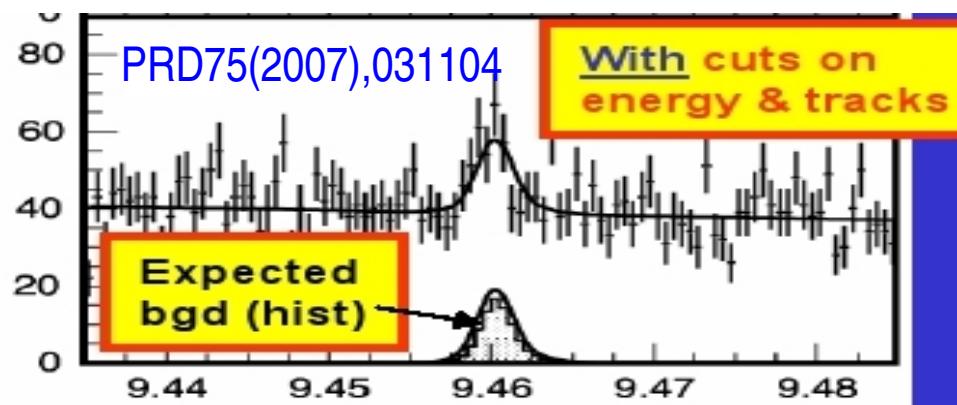
Belle: $\Upsilon(3S) \rightarrow \pi\pi + \text{NOTHING}$

$N_{\text{sig}} = 38 \pm 39(\text{stat}) \Leftrightarrow 0$ consistent

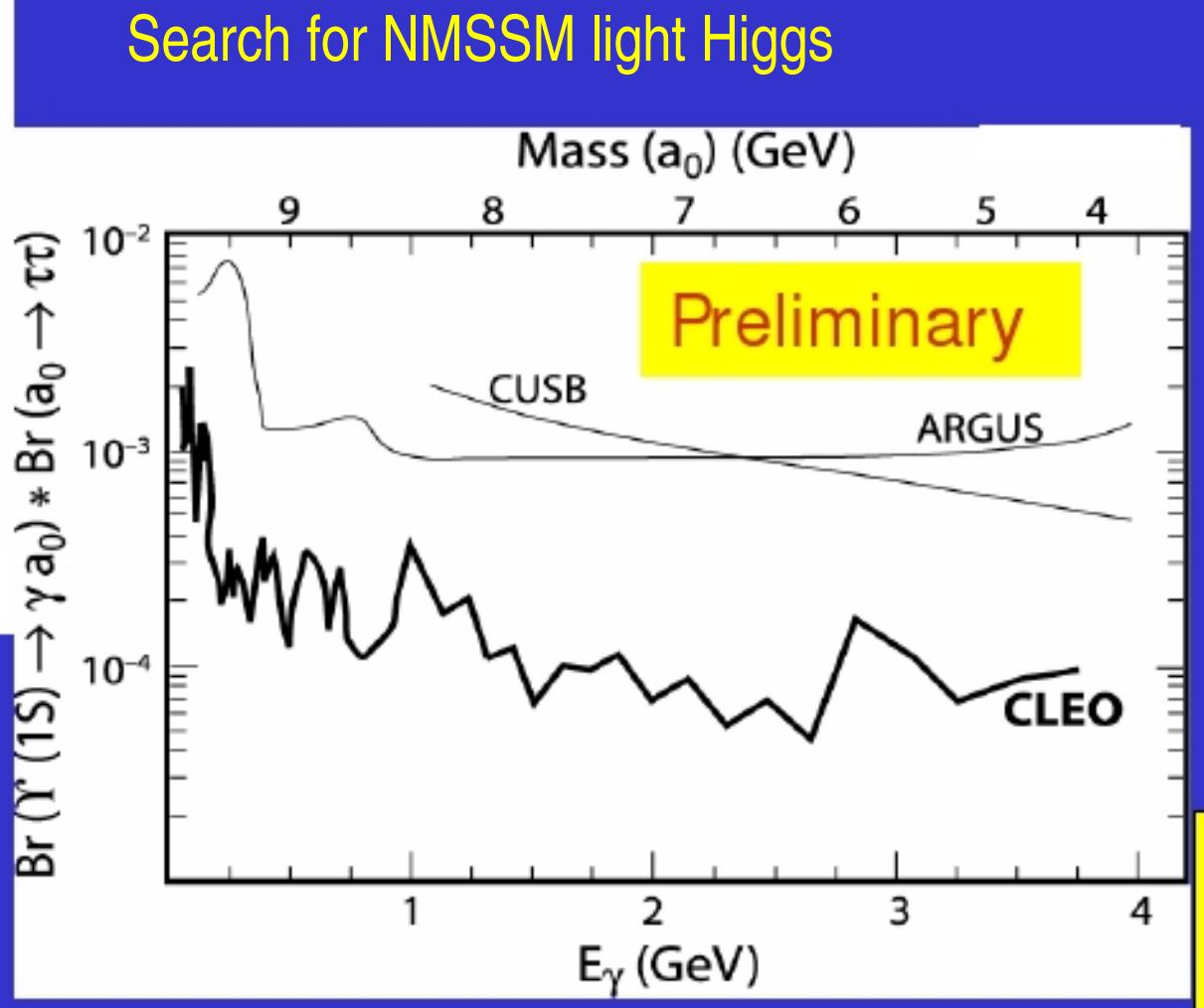
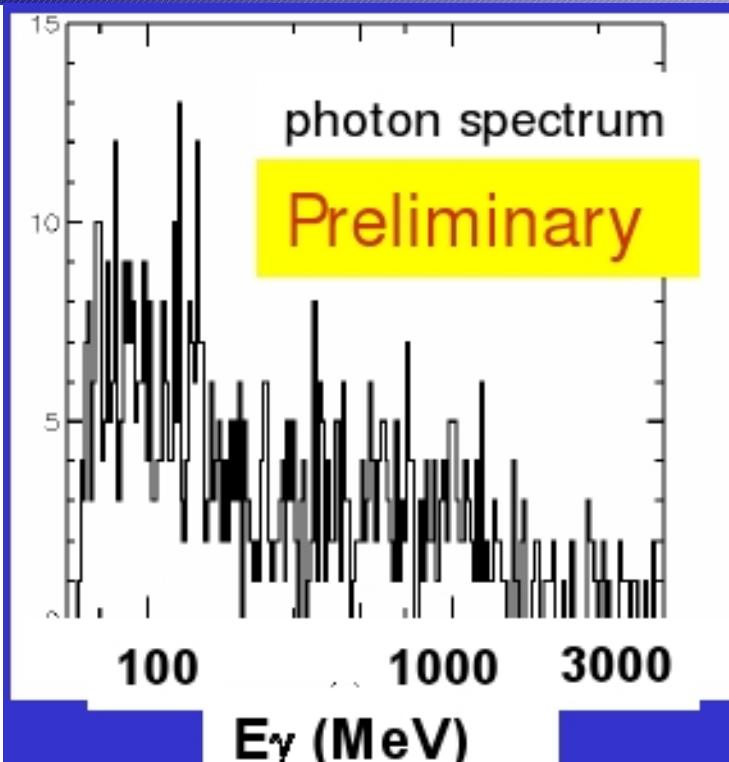


$\text{Br}(\Upsilon(1S) \rightarrow \text{invisible}) < 2.5 \times 10^{-3}$ (@90% C.L.)

Cleo: $\Upsilon(2S) \rightarrow \pi\pi + \text{NOTHING}$



New Physics searches from Y(1S) peak



Recent limits by CLEO-III , from the Y(1S) sample